

“Freeze-out” temperatures

Particle annihilation

electron + positron → gamma-ray photon + gamma-ray photon

$kT < m_{\text{particle}} c^2$

T (K)	particle / antiparticle	time since BB
6×10^{10}	e	20 sec
1.2×10^{10}	μ	1 sec
1×10^{10}	p	10

NOTE: after freeze-out, must have been an excess of matter over antimatter

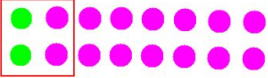
neutrinos (ν)

- $\bar{\nu} + \text{proton} \leftrightarrow e^+ + \text{neutron}$
 - $e^- + \text{proton} \leftrightarrow \text{neutron} + \nu$
- } ν energy $\sim 10^{10}$ K
- n, p, in equilibrium when $T > 10^{10}$ K
 - $\nu + n \rightarrow p^+ + e^-$ favored for $T < 10^{10}$ K
 - neutrinos “decouple”
 - $T < 10^{10}$ K - p preferred over n
 - still hot enough for $p + n \leftrightarrow \text{deuterium}$

The First Three Minutes

time	Temp [K]	Z (redshift)	Radius (?)	Density [g/cc]	Stuff
0:00:00.01	10^{11}	3×10^{10}	10 ly	4×10^9	<ul style="list-style-type: none"> • e^+, e^- • photons, neutrinos • protons(p^+) = neutrons(n) (1)
0:00:00.1	3×10^{10}	10^{10}	30 ly	3×10^7	<ul style="list-style-type: none"> • e^+, e^- • photons, neutrinos • 62% p^+, 38% n
0:00:01	10^{10}	3×10^9	100 ly	4×10^5	<ul style="list-style-type: none"> • e^+, e^- • photons • neutrinos -- decoupled! • 76% p^+, 24% n

Big Bang Nucleosynthesis

- final result as $T \sim 3 \times 10^9$ K: 87% p , 13% n
 - **10 seconds after BB!**
- 2 out of 16 baryons are neutrons
 - 
- make 1 Helium for every 12 Hydrogens
- mass fraction of Helium = $4/16 = 25\%$
- **Big Bang Nucleosynthesis should produce a universe that is 25% helium, by mass**
- predicted by G. Gamow in the 1940s - verified by observations of 1st stars

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0:00:13.8	3×10^9	10^9	300 ly	3×10^3	<ul style="list-style-type: none"> • e^- frozen out • photons, neutrinos • 83% p^+, 17% n

a complication - the “deuterium bottleneck”

- deuterium (d) = ${}^2\text{H} = n+p$
- essential step to go from $\text{H} \rightarrow {}^4\text{He}$:
 - $p + p \rightarrow d + e^+ + \gamma$
 - $p + d \rightarrow {}^3\text{He} + p \rightarrow {}^4\text{He}$
- **BUT:** $d + \gamma \leftrightarrow p + n$ easily destroys when $T > 10^9$ K
- can't finish the chain to ${}^4\text{He}$
 - deuterium doesn't hang around long enough
- $T < 10^9$ K:
 - d is stable, $d+p \rightarrow {}^3\text{He}$ ($+ p \rightarrow {}^4\text{He}$) can proceed
 - **3 minutes 52 seconds after Big Bang**

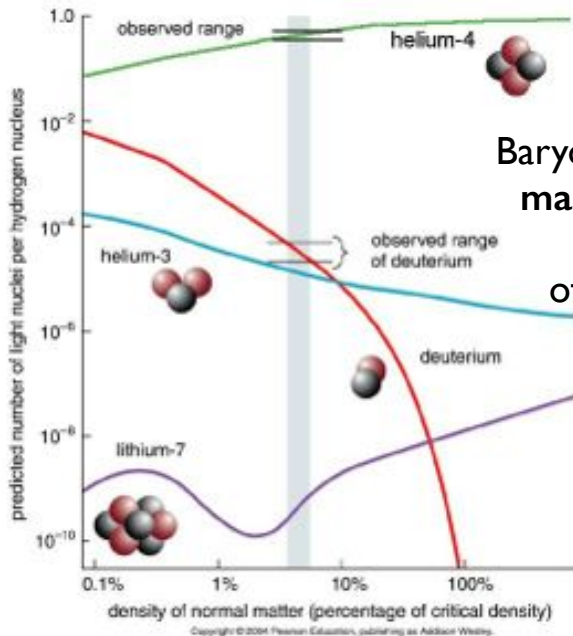
0:00:13.8	3×10^9	10^9	300 ly	3×10^3	<ul style="list-style-type: none"> e- frozen out photons, neutrinos 83% p+, 17% n
0:03:02	10^9	3×10^8	1000 ly	40	<ul style="list-style-type: none"> photons, neutrinos 86% p+, 14% n deuterium bottleneck (2)
0:03:42	9×10^8	$\sim 3 \times 10^8$	~ 1000 ly	26	<ul style="list-style-type: none"> photons, neutrinos 87% p+, 13% n (in He) deuterium now stable -> 26% Helium (3)
0:34:40	3×10^8	10^8	3000 ly	0.10	<ul style="list-style-type: none"> e+, e- photons, neutrinos 87% p+, 13% n (in He) trace ^3He, ^7Li, deuterium

Adapted from *The First Three Minutes*, by Steven Weinberg

- 1) antineutrino + proton \leftrightarrow e+ + neutron
e- + proton \leftrightarrow neutron + neutrino
- 2) proton + neutron \leftrightarrow deuterium + photon
- 3) deuterium + proton \rightarrow ^3He

more nucleosynthesis

- residual d, ^3He sensitive to baryon density
- d \uparrow , ^3He \downarrow in high density Universe
- d and ^3He probe baryonic density of early Universe
- measure cosmic d, ^3He to determine Universe density independent of other "local" measures
- ^7Li is another probe



BBN Result:
Baryonic (ordinary) matter makes up at most a few percent of the critical density of the Universe

revisiting the critical density

$$v_{\text{esc}}^2 = \frac{2GM}{R} = c^2 \quad \text{and} \quad \rho = \frac{M}{R} \frac{3}{4\pi R^2}$$

so we can write: $c^2 = 2G \times \frac{4\pi R^2 \rho}{3}$

Hubble Law: $v = H_0 \times R = c$ - so - $R = c/H_0$

$$c^2 = 2G \times \frac{4\pi c^2 \rho}{3H_0^2} \quad \text{rearrange:} \quad \rho_{\text{crit, today}} = \frac{3H_0^2}{8\pi G}$$

$$\rho_{\text{crit}} = 3H_0^2/8\pi G \sim 9.1 \times 10^{-30} \text{ g/cc} \times (H_0/70)^2$$

Is the Universe a black hole?

- define 'horizon' using age of Universe from H_0

- **critical density:**

$$\rho_{\text{crit}} = 3H^2/8\pi G \sim 9.1 \times 10^{-30} \text{ g/cc} \times (H_0/70)^2$$

- **measured density: ρ**

- recast as $\Omega_0 \equiv \rho/\rho_{\text{crit}}$

- if $\Omega < 1$: expansion **continues forever**: universe is "open"

- if $\Omega > 1$: expansion **reverses**: universe is "closed"

- **Open (infinite) Universe:**

- infinite volume no true edge

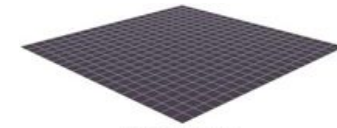
- **Closed (finite) Universe:**

- finite volume no true edge

- **Flat Universe:** density = critical density $\Omega = 1$

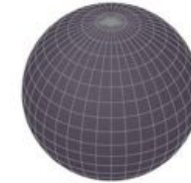
Three possible geometries

- Flat: $\Omega = 1$



flat (critical) geometry

- Closed: $\Omega > 1$



spherical (closed) geometry

- Open: $\Omega < 1$



saddle-shaped (open) geometry

The need for dark matter

- Big Bang Nucleosynthesis:

- constrained by ^3He and deuterium

- $\Omega_{\text{baryon}} < 0.1$

- There **MUST** be some gravitating stuff made of exotic, **NON-BARYONIC** matter

- see the web links...

Probing matter via dynamics

- Kepler's 3rd law gives orbital period in terms of orbital distance and mass within the orbit:

$$P_{\text{orb}}^2 \propto \frac{d^3}{M_{\text{within}}}$$

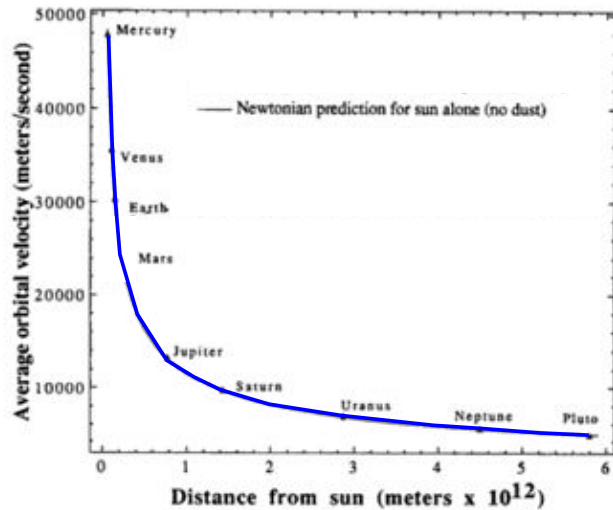
- the orbital speed is simply

$$v_{\text{orb}} = 2\pi d / P_{\text{orb}} \text{ -so-}$$

$$v_{\text{orb}} \propto \sqrt{M_{\text{within}} / d}$$

Probing matter via dynamics

$$v_{orb} \propto \sqrt{M_{within}/d}$$



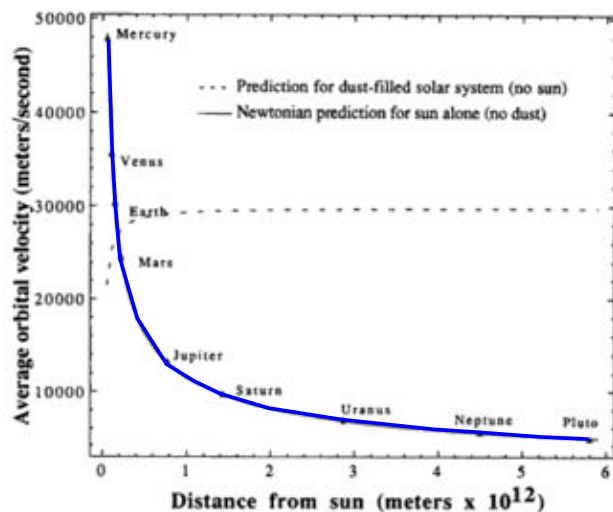
from 'Qunitessence' by Lawrence Krauss

Probing matter via dynamics

- what if... solar system was filled with invisible dust?
- can put $\sim 1M_{\odot}$ of dust within the Earth's orbit at very low density (10^{-4} x air)
- if evenly distributed
 - M within orbits would *increase* as d increases
 - expect higher V than Keplerian
 - faster orbital speed than without dust:

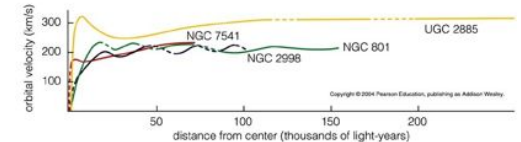
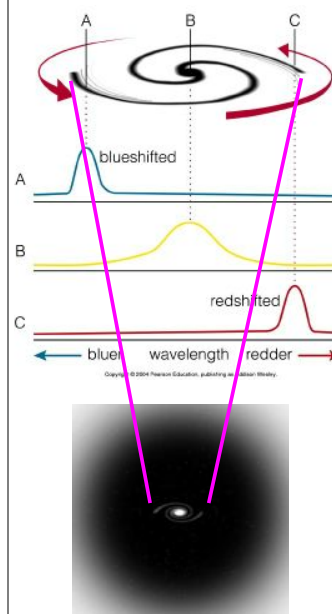
Probing matter via dynamics

$$v_{orb} \propto \sqrt{M_{within}/d}$$



from 'Qunitessence' by Lawrence Krauss

Dark Matter (within galaxies)



- "Rotation curve" gives rotation speed via spectroscopy
- Rotation rate of stars/clouds gives mass within the orbit
- **Spiral Galaxies** - matter extends well beyond visible disks
- **Elliptical Galaxies** - big M/L ratio
- **90% of mass in galaxies is dark**

mass-density of the Universe (in galaxies)

- measure "average" mass of a galaxy via rotation curve
- count the number of galaxies
- determine **mass** (in galaxies) of the universe
- divide by **volume** of the galaxy sample to get Ω

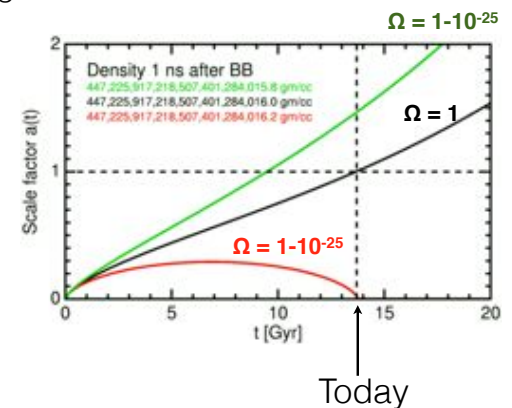


so what if $\Omega_0 < 1$?

- if $\Omega_0 > 0.1$, then the dark matter in galaxies must be **non-baryonic (!?)**
- the "flatness problem" looms:
 - if $\Omega_0 = 1.0$ now, then Ω has *always* been 1.0
 - if $\Omega_0 < 1.0$ now, it was much closer to 1.0 in the distant past, but still less than 1.0
 - if $\Omega_0 > 1.0$ now, it was much closer to 1.0 in the distant past, but still greater than 1.0

the flatness problem

- if not *exactly* 1.000... at the Big Bang, then Ω diverges very rapidly away from it as the universe expands.
- If Ω was only slightly *smaller* than 1.000... at the Big Bang, then Ω_0 should be *nearly zero* today.
- If Ω was only slightly *bigger* than 1.000... at the Big Bang, the Universe should have collapsed long ago.



(From Ned Wright's cosmology tutorial)